

This is a repository copy of *Progress in understanding disruptions triggered by massive gas injection via 3D non-linear MHD modelling with JOEUK*.

White Rose Research Online URL for this paper:

<https://eprints.whiterose.ac.uk/133796/>

Version: Accepted Version

Article:

Nardon, E., Fil, A., Hoelzl, M. et al. (1 more author) (2017) Progress in understanding disruptions triggered by massive gas injection via 3D non-linear MHD modelling with JOEUK. Plasma Physics and Controlled Fusion. 014006. ISSN 1361-6587

<https://doi.org/10.1088/0741-3335/59/1/014006>

Reuse

Items deposited in White Rose Research Online are protected by copyright, with all rights reserved unless indicated otherwise. They may be downloaded and/or printed for private study, or other acts as permitted by national copyright laws. The publisher or other rights holders may allow further reproduction and re-use of the full text version. This is indicated by the licence information on the White Rose Research Online record for the item.

Takedown

If you consider content in White Rose Research Online to be in breach of UK law, please notify us by emailing eprints@whiterose.ac.uk including the URL of the record and the reason for the withdrawal request.

Progress in understanding disruptions triggered by massive gas injection via 3D non-linear MHD modelling with JOREK

E. Nardon¹, A. Fil¹‡, M. Hoelzl², G. Huijsmans¹ and JET contributors³§

¹CEA, IRFM, F-13108 Saint-Paul-lez-Durance, France

²Max Planck Institute for Plasma Physics, Boltzmannstr. 2, 85748 Garching b. M., Germany

³EUROfusion Consortium, JET, Culham Science Centre, Abingdon, OX14 3DB, UK

Abstract.

3D non-linear MHD simulations of a D_2 Massive Gas Injection (MGI) triggered disruption in JET with the JOREK code provide results which are qualitatively consistent with experimental observations and shed light on the physics at play. In particular, it is observed that the gas destabilizes a large $m/n = 2/1$ tearing mode, with the island O-point coinciding with the gas deposition region, by enhancing the plasma resistivity via cooling. When the $2/1$ island gets so large that its inner side reaches the $q = 3/2$ surface, a $3/2$ tearing mode grows. Simulations suggest that this is due to a steepening of the current profile right inside $q = 3/2$. Magnetic field stochastization over a large fraction of the minor radius as well as the growth of higher n modes ensue rapidly, leading to the Thermal Quench (TQ). The role of the $1/1$ internal kink mode is discussed. An I_p spike at the TQ is obtained in the simulations but with a smaller amplitude than in the experiment. Possible reasons are discussed.

‡ Present address: Princeton University, NJ 08540, US

§ See the Appendix of F. Romanelli et al., Proceedings of the 25th IAEA Fusion Energy Conference 2014, Saint Petersburg, Russia

1. Introduction

In order to develop a predictive modelling capability in support of the design and operation of the ITER Disruption Mitigation System (DMS) [1] [2], it is essential to validate models on present devices as well as to understand the physical mechanisms at play. The present paper describes progress made with the JOE2 3D non-linear magnetohydrodynamics (MHD) code in the domain of Massive Gas Injection (MGI), one of the options considered for the ITER DMS. The focus is put on the pre-Thermal Quench (TQ) and TQ phases. The Current Quench (CQ) phase will be the object of future work.

A publication by Fil et al. in 2015 [3] introduced first results from simulations of a pure D_2 MGI into a JET Ohmic plasma. Work has been pursued since then on the same case and with essentially the same model, but with more realistic input parameters thanks to numerical improvements. This has led to significant progress which is presented in this paper. Most notably, while only an incomplete TQ was obtained in [3], simulations now display a more proper TQ. Furthermore, numerical experiments as well as a detailed analysis of simulation results have led to a better understanding of the physics at play.

The paper is constructed as follows. In Section 2, the experiment is introduced. Section 3 presents the model and Section 4 the setup of simulation parameters, highlighting differences with [3]. Section 5 gives an overview of a simulation, after which Sections 6 and 7 describe investigations on the mechanisms at play during the pre-TQ and TQ phases, respectively. Finally, Section 8 concludes.

2. The experiment

The modelled case has already been described in the paper by Fil et al. [3] but it is briefly described again here for convenience. This is JET pulse 86887, an Ohmic pulse with $B_t = 2$ T, $I_p = 2$ MA, $q_{95} = 2.9$ in which a disruption was MGI-triggered on a “healthy” plasma by activating the Disruption Mitigation Valve number 2 (DMV2) pre-loaded with D_2 at 5 bar. The gas is injected at the outer midplane. Central values of the electron density and temperature prior to the MGI are $n_e = 3 \cdot 10^{19} \text{ m}^{-3}$ and $T_e = 1.2$ keV. The volume of the DMV2 reservoir is 10^{-3} m^3 and its temperature is about 300 K, so it initially contains about $1.2 \cdot 10^{23}$ D_2 molecules, which represents roughly 100 times more D nuclei than initially present in the plasma.

Figure 1 shows an overview of the disruption phase. First effects of the MGI are visible from about 2 ms (relative to the DMV2 trigger) in the form of increases in the line integrated density and radiated power. The TQ occurs at about 12 ms as can be seen

from the fast collapse of the central Soft X-Ray (SXR) signal accompanied by a burst of MHD activity and immediately followed by the characteristic I_p spike. The CQ ensues.

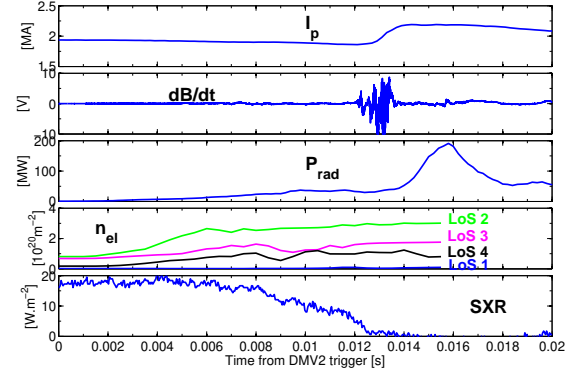


Figure 1. Experimental time traces for JET pulse 86887, from top to bottom: plasma current I_p , magnetic fluctuations from Mirnov coil, radiated power from bolometry, line integrated density from interferometry, and soft X-rays signal from a central chord. The time origin corresponds to the DMV2 trigger.

3. The model

The model used here is essentially the same as in [3] (with only a few differences detailed below). The reader is therefore referred to [3] for a more complete model description. Nevertheless, for convenience we write again the equations here:

$$\frac{\partial \psi}{\partial t} = \eta \Delta^* \psi - R[u, \psi] - F_0 \frac{\partial u}{\partial \phi} \quad (1)$$

$$j = \Delta^* \psi \quad (2)$$

$$\begin{aligned} R \nabla \cdot \left(R^2 \rho \nabla_{pol} \frac{\partial u}{\partial t} \right) &= \frac{1}{2} \left[R^2 |\nabla_{pol} u|^2, R^2 \rho \right] + [R^4 \rho \omega, u] \\ &+ [\psi, j] - \frac{F_0}{R} \frac{\partial j}{\partial \phi} + [\rho T, R^2] + R \mu \nabla^2 \omega \\ &+ \nabla \cdot ((\rho \rho_n S_{ion} - \rho^2 \alpha_{rec}) R^2 \nabla_{pol} u) \end{aligned} \quad (3)$$

$$\omega = \nabla_{pol}^2 u = \frac{1}{R} \frac{d}{dR} \left(R \frac{du}{dR} \right) + \frac{d^2 u}{dZ^2} \quad (4)$$

$$\frac{\partial \rho}{\partial t} = -\nabla \cdot (\rho \mathbf{v}) + \nabla \cdot (D_{\perp} \nabla_{\perp} \rho + D_{\parallel} \nabla_{\parallel} \rho) + \rho \rho_n S_{ion} - \rho^2 \alpha_{rec} \quad (5)$$

$$\begin{aligned} \frac{\partial(\rho T)}{\partial t} = & -\mathbf{v} \cdot \nabla(\rho T) - \gamma \rho T \nabla \cdot \mathbf{v} \\ & + \nabla \cdot (\kappa_{\perp} \nabla_{\perp} T + \kappa_{\parallel} \nabla_{\parallel} T) \\ & + \frac{2}{3R^2} \eta j^2 - \xi_{ion} \rho \rho_n S_{ion} \end{aligned} \quad (6)$$

$$\begin{aligned} \rho B^2 \frac{\partial v_{\parallel}}{\partial t} = & -\rho \frac{F_0}{2R^2} \frac{\partial(B^2 v_{\parallel}^2)}{\partial \phi} - \frac{\rho}{2R} [B^2 v_{\parallel}^2, \psi] \\ & - \frac{F_0}{R^2} \frac{\partial(\rho T)}{\partial \phi} + \frac{1}{R} [\psi, \rho T] + B^2 \mu_{\parallel}(T) \nabla_{pol}^2 v_{\parallel} \\ & + (\rho^2 \alpha_{rec} - \rho \rho_n S_{ion}) B^2 v_{\parallel} \end{aligned} \quad (7)$$

$$\frac{\partial \rho_n}{\partial t} = \nabla \cdot (D_n : \nabla \rho_n) - \rho \rho_n S_{ion} + \rho^2 \alpha_{rec} + S_n \quad (8)$$

ψ is the poloidal flux, $j = R j_{\phi}$ the toroidal current density times the major radius, u the poloidal flow potential, ω the toroidal vorticity, ρ the plasma mass density, T the total (ion + electron) temperature, v_{\parallel} the parallel (to the magnetic field) velocity and ρ_n the neutral mass density. S_{ion} and α_{rec} designate respectively the ionization and recombination rate coefficients for deuterium, parameterized as detailed in [3]. Note that apart from the neutral source, there are no sources in the model. Realistic sources indeed have a negligible effect on the small timescale of the disruption. Hyper-resistivity/conductivity/viscosity terms, not shown above, are also included for numerical stability reasons.

There are two differences between the present model and that of [3]. First, for the sake of energy conservation, the resistivity used in the Ohmic heating term of Eq. 6 is the same as in Eq. 1, while in [3] the Spitzer resistivity was used. Second, D radiation terms have been suppressed for simplicity after verifying that they do not influence the results significantly.

4. Simulations setup

4.1. Initial conditions and general parameters

Initial density and temperature profiles are taken from Thomson scattering measurements, assuming $T_i = T_e$. The initial magnetic equilibrium is calculated from EFIT data as explained in [3]. Toroidal rotation is initialized at 0. Table 1 gives the value of input parameters used here as well as in [3]. Improvements in the poloidal mesh now allow running simulations with a hyper-resistivity η_h and hyper-viscosity μ_h roughly 500 times smaller than in [3], which removes the artificial stabilization of the MHD activity due to these coefficients in [3]. Also, present simulations use much smaller diffusion coefficients than in [3],

both for ions/electrons and for neutrals. The artificial core dilution observed in [3] is therefore not present anymore. It is important to recall that realistic temperature dependencies are used for the resistivity and parallel thermal conductivity: $\eta = \eta_0 \cdot (T_0/T)^{3/2}$ and $\kappa_{\parallel} = \kappa_{\parallel 0} \cdot (T/T_0)^{5/2}$, where T_0 is the initial temperature at the center of the plasma. For $\kappa_{\parallel 0}$, the Spitzer-Härm value is used, while η_0 is typically one order of magnitude larger than the Spitzer value.

4.2. Gas source setting

The physics of gas penetration into the plasma during an MGI is rather complex [4] and it is clear that the present model is not appropriate to describe it accurately. In other terms, JOE3 simulations are not predictive in this respect. The gas source is therefore adjusted so as to best match interferometry data. Its temporal shape is based on the solution for the gas flow in vacuum [3] but using a DMV2 pressure of 0.1 bar instead of the experimental 5 bar (which can be interpreted as a sign of the low fuelling efficiency of the MGI). Note that in [3], a much larger value of typically 1 bar was needed due to the use of much larger particle diffusion coefficients. The gas source is located at the edge of the plasma (centered 5 cm inside the separatrix) and localised both poloidally and toroidally (see [3] for precisions). Figure 2 shows experimental and synthetic line integrated densities n_{el} for three interferometry Lines of Sight (LoS). Three simulations are shown, with different toroidal localisations of the source: $\Delta\phi_{MGI} = 0.5$ (dashed), 2 (plain) and 100 (i.e. virtually axisymmetric, dash-dotted) respectively (angles are given in radians). Even though the same number of neutrals is injected in each simulation, clear differences appear in the synthetic n_{el} . In order to interpret these differences, one should consider that interferometry chords are located 180 degrees away toroidally from DMV2, as shown in Fig. 3. Poloidal cross-sections of n_e in the interferometer plane are shown in Fig. 4 for $\Delta\phi_{MGI} = 0.5$ (top) and $\Delta\phi_{MGI} = 100$ (bottom). These figures indicate that LoS 2 (red) runs through a region which is connected to the gas deposition region and therefore sees a faster increase in n_{el} than LoS 3 (blue) for $\Delta\phi_{MGI} = 0.5$ (dashed) but not for $\Delta\phi_{MGI} = 100$ (dash-dotted). The simulation at $\Delta\phi_{MGI} = 0.5$ actually matches experimental data very well for the three LoS until about 3 ms, after which the agreement deteriorates for a yet unidentified reason. Over a larger period of time, the case $\Delta\phi_{MGI} = 2$ is the one that offers the best global match and is therefore selected for a detailed description and analysis below.

Parameter	Symbol	Value in present paper	Value from [3]	Units	Comment
Initial central resistivity	η_0	$3.5 \times 10^{-8} - 3.5 \times 10^{-7}$	Same	SI ($\Omega \cdot \text{m}$)	Spitzer value: 2×10^{-8}
Resistive time at $q = 2$	$\tau_R^{q=2}$	$4 - 0.4$	Same	SI (s)	Exp. value: 7
Lundquist number at $q = 2$	$S^{q=2} = \tau_R^{q=2} / \tau_A^{q=2}$	$3.4 \times 10^7 - 3.4 \times 10^6$	Same	Adimensional	Exp. value: 6×10^7
Hyper-resistivity	η_h	1×10^{-12}	5×10^{-10}	JOREK	
Initial central // thermal conductivity	$\kappa_{\parallel 0}$	3×10^{29}	7×10^{28}	SI ($/\text{m/s}$)	Spitzer-Härm value: 3×10^{29}
\perp thermal conductivity	κ_{\perp}	4×10^{19}	Same	SI ($/\text{m/s}$)	Corresponds to $\chi_{\perp} \simeq 1 \text{ m}^2/\text{s}$
Hyper- \perp thermal conductivity	$\kappa_{\perp h}$	10^{-12}	Same	JOREK	
Dynamic viscosity	μ	1.4×10^{-7}	2.8×10^{-7}	SI (kg/m/s)	Corresponds to $\nu \simeq 1 \text{ m}^2/\text{s}$
Hyper-viscosity	μ_h	1×10^{-12}	4×10^{-10}	JOREK	
// viscosity	μ_{\parallel}	2.8×10^{-6}	2.8×10^{-5}	SI (kg/m/s)	Corresponds to $\nu \simeq 20 \text{ m}^2/\text{s}$
Hyper-// viscosity	$\mu_{\parallel h}$	10^{-9}	Same	JOREK	
Ion/electron particle \perp diffusivity	D_{\perp}	1.4	28	SI (m^2/s)	
Ion/electron particle // diffusivity	D_{\parallel}	0	2.8×10^4	SI (m^2/s)	
Neutral particle diffusivity	D_n	60 – 6000	2.8×10^4	SI (m^2/s)	

Table 1. List of typical values of JOREK input parameters used in the present paper as well as in [3]

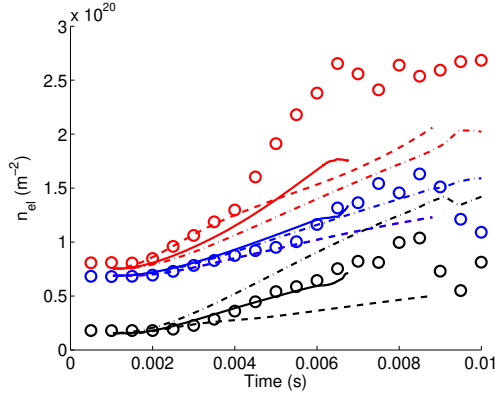


Figure 2. Time evolution of the measured (circles) and simulated (lines - plain = $\Delta\phi_{MGI} = 2$; dashed = $\Delta\phi_{MGI} = 0.5$; dash-dotted = $\Delta\phi_{MGI} = 100$) line integrated density for the three interferometry lines of sight

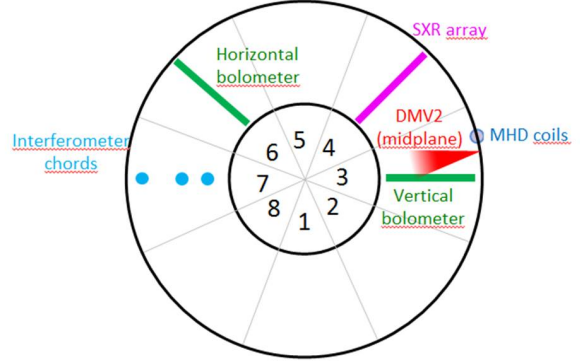


Figure 3. Schematic machine view from the top showing the DMV2 and interferometer chords

5. Overview of a simulation

Snapshots of the above simulation with $\Delta\phi_{MGI} = 2$ are shown in Figure 5. For a series of times through the simulation (different rows), poloidal cross sections of (from left to right) T_e , n_e , j_ϕ as well as a Poincaré plot are shown. The top row shows the initial state. The second row shows the situation at $t = 4.1$ ms, during the pre-TQ phase. It can be seen that the gas has increased n_e and decreased T_e at the edge of the plasma and given birth to an $m/n = 2/1$ magnetic island, clearly visible on the Poincaré plot and whose effect on T_e , n_e and j_ϕ is also clear. The island O-point coincides with the gas deposition region, which is in line with experimental observations for Ohmic plasmas in JET [5] (we will come back to this point

in Section 6). A $1/1$ island is also visible in the core which indicates the growth of an internal kink mode ($q_0 < 1$ in this simulation, consistently with the presence of sawteeth in this pulse). In the third row, at $t = 5.7$ ms, the magnetic field has become stochastic over roughly the outer half of the plasma, which leads to a temperature flattening by parallel conduction in this region. However, the core is still hot at this stage. Finally, at $t = 6.2$ ms (fourth row), the core T_e has collapsed via a convective mixing related to the $1/1$ mode. The behaviour observed in this simulation is similar to the one found in NIMROD simulations for other tokamaks [6] [7].

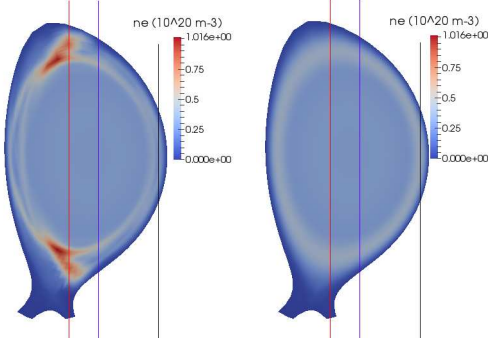


Figure 4. Poloidal cross-sections of n_e in the plane of the interferometer as simulated with JOREK for $\Delta\phi_{MGI} = 0.5$ (left) and $\Delta\phi_{MGI} = 100$ (right). Interferometry chords are also shown, with the same color code as in Fig. 2.

6. Physics of the pre-thermal quench phase

Let us now analyse the mechanisms at play, beginning with the pre-TQ phase. An important question is how the gas generates the 2/1 tearing mode. One may imagine at least three possible mechanisms. The first one is that the gas creates a 3D pressure field to which \mathbf{j} and \mathbf{B} have to adjust in order for force balance to pertain, leading to a 3D equilibrium which may imply a 2/1 tearing mode. The other two mechanisms, in contrast to the first one, are directly related to the temperature dependence of the resistivity η . One is an axisymmetric effect: the gas cools the edge of the plasma, increasing η and therefore contracting the current profile. This should increase the current density gradient $|j'_\phi|$ at the edge of the still hot region. When the steep gradient gets just inside the $q = 2$ surface, the 2/1 tearing mode should be destabilized via the following term in the resistive MHD energy principle [8]: $\delta W \propto \int \frac{j'_\phi}{1-nq/m} |\psi_1|^2 r dr$. The other η -related effect is a non-axisymmetric one: the gas cools the plasma locally and therefore suppresses the current locally, which should lead to the growth of magnetic islands with their O-point at the gas deposition location. This mechanism is involved as well in radiation driven islands, which have been invoked recently to explain density limit disruptions [9]. To a certain extent, numerical experiments allow discriminating between the three above-mentioned mechanisms. Figure 6 shows the evolution of the magnetic energy contained in the $n = 1$ toroidal harmonic for three simulations. In the first one (blue), the standard $\eta = \eta(T)$ dependence was used. In the second one (red), we used instead $\eta = \eta(T_{n=0})$, where $T_{n=0}$ is the axisymmetric component of T , in order to suppress the non-axisymmetric η -related mechanism. Finally, in the third simulation (black), the T dependence of η was removed altogether

by keeping the initial η profile, in order to remove both η -related effects. All three simulations are overimposed until about 0.5 ms, indicating that the early evolution is not η -related and therefore likely connected to the 3D equilibrium mechanism. However, in the third (black) simulation the $n = 1$ magnetic energy quickly saturates and the 2/1 island remains very small. The leading mechanisms for the 2/1 island growth are therefore η -related, even if the island grows from a non- η -related seed. Then, it appears that between 0.5 and 2.5 ms, the $\eta = \eta(T)$ case gains about one order of magnitude compared to the $\eta = \eta(T_{n=0})$ case, showing that the non-axisymmetric η -related effect is important. This is consistent with the already mentioned experimental observation that the island O-point coincides with the gas deposition region for Ohmic plasmas [5]. Note that for neutral beam heated plasmas, the alignment is observed to degrade [5], which may be due to plasma rotation via two effects: a smearing of the non-axisymmetric η -related effect and a drag of the island into the plasma flow.

7. Physics of the thermal quench

7.1. Triggering of the thermal quench

The next question is: what mechanisms trigger the TQ? This question may be addressed by analysing Figures 7 and 8. Fig. 7 shows the time evolution of the magnetic energy in the different toroidal harmonics (note the logarithmic vertical scale). Fig. 8 contains poloidal cross-sections of the cosine component of ψ for the $n = 1$ (left), $n = 2$ (middle) and $n = 3$ (right) harmonics at different times (different rows). Both figures are for the same simulation as in Section 5. It can be seen in Fig. 7 that during the first 5 ms, magnetic perturbations are strongly dominated by the $n = 1$ mode. The top left plot of Fig. 8 shows that the mode has a 2/1 structure, not surprisingly. Then, at about 5 ms, a clear increase in the growth rate of the $n = 2$ magnetic energy is visible. This is associated to the growth of a 3/2 tearing mode, as can be seen by comparing the first and second rows, middle column, plots in Fig. 8. Around 5.8 ms, magnetic energies in higher n harmonics grow sharply. For example, a 4/3 mode can be seen to grow when comparing the second and third rows, last column, plots in Fig. 8. This leads to the TQ.

Mechanisms for the growth of the 2/1 mode have been discussed in Section 6. What about the other modes? Looking at Poincaré plots, it appears that the 3/2 mode starts growing when the inner side of the 2/1 island reaches the $q = 3/2$ surface. Observing the $j_{\phi,n=0}$ profile at different times, shown in Figure 9, it seems likely that this is due to a current profile effect. Indeed, the 2/1 island causes a local $j_{\phi,n=0}$ flattening,

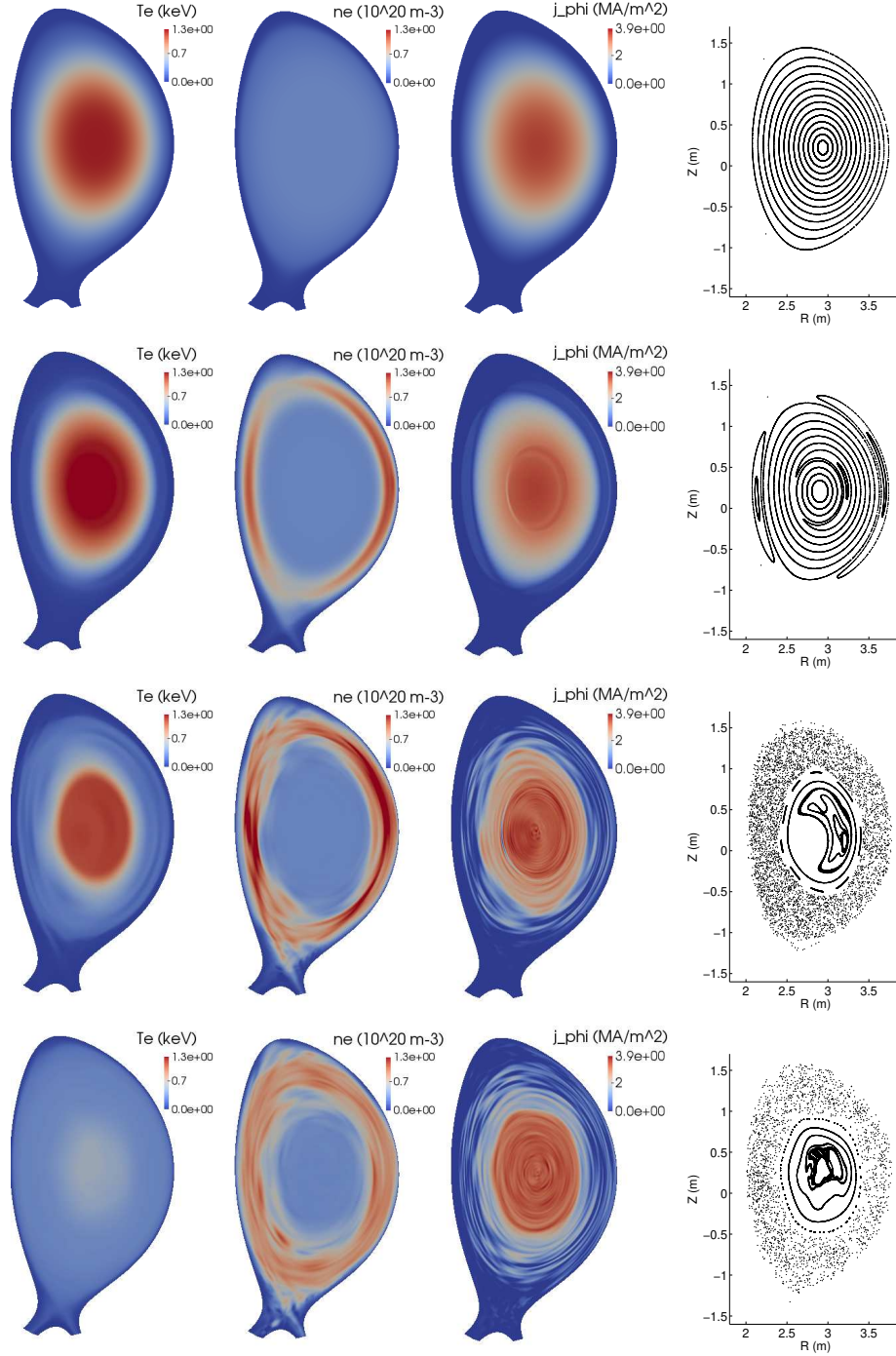


Figure 5. Poloidal cross sections at the toroidal position of the MGI of (from left to right) T_e , n_e , j_ϕ and Poincaré plots at times (from top to bottom) $t = 0$, $t = 4.1$ ms, $t = 5.7$ ms and $t = 6.2$ ms

which steepens the profile radially inward (compare the black, blue and red profiles). The $3/2$ mode grows when the large $|j'_{\phi,n=0}|$ region runs across the $q = 3/2$ surface (i.e. between the blue and red profiles) due to the δW term in the energy principle already mentioned in Section 6. The same process then seems to take place with the $3/2$ tearing mode, which locally flattens $j_{\phi,n=0}$

(red vs. magenta profiles), moving the large $|j'_{\phi,n=0}|$ region across $q = 4/3$ and destabilizing the $4/3$ mode. Therefore, it seems that the TQ is triggered by a kind of current profile avalanche. Note that this picture for the TQ triggering has already been proposed in [6] and [8]. The magnetic stochasticity created by the tearing modes in (roughly) the outer half of the

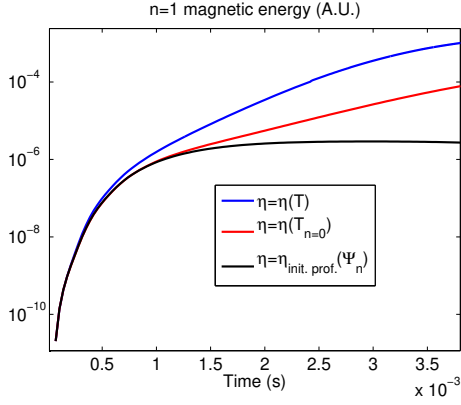


Figure 6. Time evolution of the magnetic energy contained in the $n = 1$ component for three simulations with different models for the resistivity (see text)

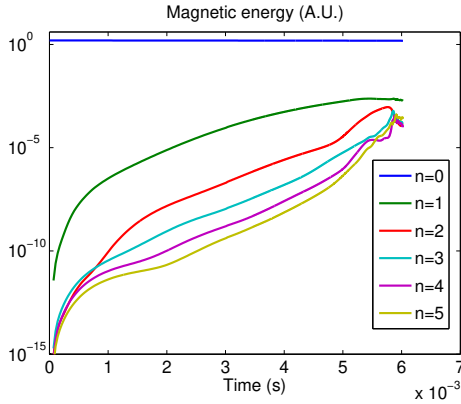


Figure 7. Time evolution of the magnetic energy in the different toroidal harmonics

plasma flattens the temperature profile there. In this simulation, flux surfaces however pertain inside mid-radius so that the central temperature cannot collapse purely from parallel thermal conduction. Convective core mixing indeed seems to play an important role in this simulation. However, as we shall see below, it is likely that tearing modes are too weak in this simulation. In other simulations where the tearing modes were excited more strongly (due to different MGI settings for example), magnetic stochasticity extended across the whole plasma. Hence, convective core mixing and the 1/1 internal kink mode may not play a large role in reality, but this is not clear at this stage.

7.2. The plasma current spike

The I_p spike is a characteristic feature of disruptions. A classic explanation of its origin is that the TQ releases magnetic energy $\simeq l_i I_p^2$ (where l_i is the internal inductance of the plasma) while the flux ψ at the edge of the plasma does not have the time to change

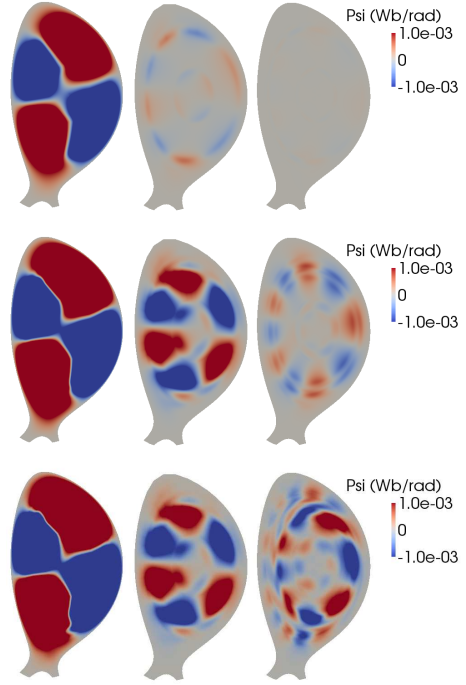


Figure 8. Poloidal cross-sections at the toroidal position of the MGI of (from left to right) the $n = 1$, $n = 2$ and $n = 3$ cosine component of ψ at times (from top to bottom) $t = 4.1$ ms, $t = 5.1$ ms and $t = 5.7$ ms (same simulation as in Figs. 5, 7 and 9). The color scale is the same for all plots (note the saturation for the $n = 1$ mode which has a large amplitude compared to the other modes).

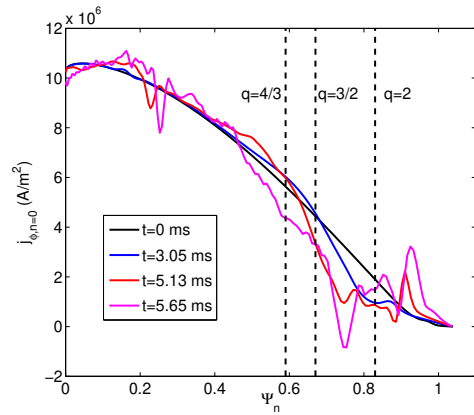


Figure 9. Toroidal current density profiles at the midplane (low field side) at different times. Note that the red and magenta profiles correspond to the last two rows of Figure 8.

significantly because the TQ is much shorter than the wall penetration time, and hence $L_p I_p$ should remain constant, where $L_p \simeq \mu_0 R [\ln(8R/a) - 2 + l_i/2]$ is the self-inductance of the plasma. The consequence is that l_i has to decrease and I_p increase, hence the I_p spike.

JOREK results are well in line with this explanation, as seen in Figure 10, which displays the evolution of I_p , l_i , the magnetic energy inside the plasma (E_{mag} and its approximation $l_i I_p^2$) and $L_p I_p$ (for a different simulation from above because the latter has a smaller I_p spike). The fact that simulations display an I_p spike in spite of having no sources answers a criticism made by Zakharov to NIMROD simulations [10] [11]. The simulated I_p spike is however much smaller than the experimental one (compare Figs. 10 and 1). Following the above argument, this should mean that not enough magnetic energy is released during the TQ and that l_i does not decrease enough, i.e. that the current profile avalanche is not sufficiently pronounced. The strength of the avalanche may be expected to have a positive dependence on the ratio between the resistive diffusion time and the growth time of the 2/1 mode: indeed, the larger this ratio, the larger the current gradient inside the 2/1 island, and the stronger the excitation of inner modes. A stronger local cooling near $q = 2$ should go in this direction. The above simulation may therefore have a too weak cooling. This is also consistent with the fact that simulations over-predict the pre-TQ duration when approaching a realistic η value. It is plausible that background impurities such as Argon remaining from previous pulses (where Ar MGI was used) [12] provide this extra cooling. The need for background impurities to match experimental observations had already been identified for He MGI simulations in Alcator C-Mod and DIII-D with NIMROD [7].

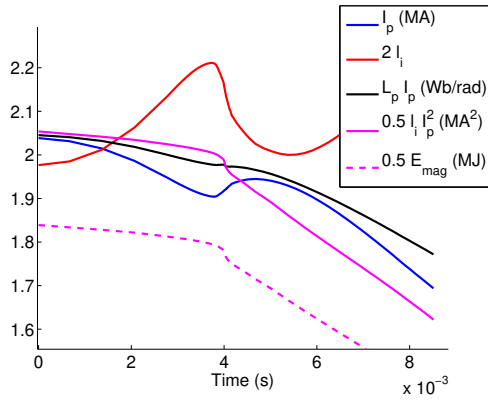


Figure 10. Time traces (from a different simulation than Figs. 5, 7, 9 and 8, in which the I_p spike is smaller) showing an I_p spike at the TQ (blue curve) with $L_p I_p$ remaining approximately constant (black curve) while the current profile flattens (red curve) and magnetic energy is dissipated (magenta curves).

8. Conclusion

JOREK simulations shed light on the physics of the pre-TQ and TQ phases of a D_2 MGI-triggered disruption in an Ohmic JET plasma. The gas destabilizes (essentially via increasing η by cooling) a 2/1 tearing mode which flattens the current profile, destabilizing a 3/2 mode, which in turn destabilizes higher n modes. The energy is lost by parallel conduction along stochastic field lines and possibly also by a convective mixing of the core, but this will have to be clarified in the future. Indeed, the I_p spike in the simulations is too weak, suggesting that tearing modes are not large enough, and simulations with larger tearing modes (which display a larger I_p spike) produce a full stochasticization of the magnetic field so that the convective core mixing plays a less important role. An interesting question which could be the object of future work is whether the TQ triggering picture found here may explain the experimental finding that the TQ is triggered at a distinct locked mode amplitude [13]. The work with JOREK on D_2 MGI simulations shall in any case be pursued, aiming for quantitative validation. For this purpose, the effect of background impurities (as well as toroidal rotation, diamagnetic effects, etc.) shall be investigated and efforts made to approach realistic parameter values, in particular for η . D_2 MGI ASDEX Upgrade simulations have been begun, for which it is easier to run at realistic parameter values. In parallel to this work, a model for MGI of other gases than D_2 has been implemented in JOREK and shall be utilized in the near future.

9. Acknowledgements

This work has been carried out within the framework of the EUROfusion Consortium and has received funding from the Euratom research and training programme 2014-2018 under grant agreement No 633053. The views and opinions expressed herein do not necessarily reflect those of the European Commission. This work was granted access to the HPC resources of Aix-Marseille University financed by the project Equip@Meso (ANR-10-EQPX-29-01). A part of this work was carried out using the HELIOS supercomputer system (IFERC-CSC), Aomori, Japan, under the Broader Approach collaboration, implemented by Fusion for Energy and JAEA, and using the CURIE supercomputer, operated into the TGCC by CEA, France, in the framework of GENCI and PRACE projects.

- [1] M. Lehnen et al., J. Nucl. Mater. **463** 39 (2015)
- [2] E. M. Hollmann et al., Phys. Plasmas **22** 021802 (2015)
- [3] A. Fil et al., Phys. Plasmas **22** 062509 (2015)
- [4] E. Nardon et al., submitted to Nucl. Fusion

- [5] M. Lehnen et al., Nucl. Fusion **55** 123027 (2015)
- [6] V.A. Izzo, Nucl. Fusion **46**, 541 (2006)
- [7] V. A. Izzo et al., Phys. Plasmas **15**, 056109 (2008)
- [8] D. Biskamp, Nonlinear Magnetohydrodynamics, Cambridge University Press, Cambridge, U.K.
- [9] D. A. Gates and L. Delgado-Aparicio, Phys. Rev. Lett. **108** 165004 (2012)
- [10] L.E. Zakharov, Nucl. Fusion **50** 058002 (2010)
- [11] V.A. Izzo and P.B. Parks, Nucl. Fusion **50** 058001 (2010)
- [12] I. Coffey, private discussion
- [13] P.C. de Vries et al., Nucl. Fusion **56** 026007 (2016)